Introduction to Basic Image Frame Calibration*

CCD, Video & DSLR

* Also known as reduction

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INTRODUCTION

- Amateur astronomy has entered the digital
 « universe » (CCD, Video, DSLR)
 - CCD and DSLR are imaging cameras ...
 - capture images for later calibration and processing
 - Video cameras are designed for digital observing ...
 - « + » image frame capture
 - « + » simultaneous calibration and processing on the fly

INTRODUCTION

- Basic image frame calibration (reduction)
 - Principles
 - Worked example for CCD
 - Principal challenges in Video
 - Evolving world of DSLR

(Sub)Frame vs Image ...

Calibration

- What is it?
 - Cleaning up « errors » and noise
- Why do it?
 - Optical « noise »
 - Dust & Dirt
 - Uneven illumination
 - Vignetting
 - Electronic or thermal noise
 - Hot pixels
 - Amplifier glow

Dust & Dirt



Vignetting



Amplifier Glow



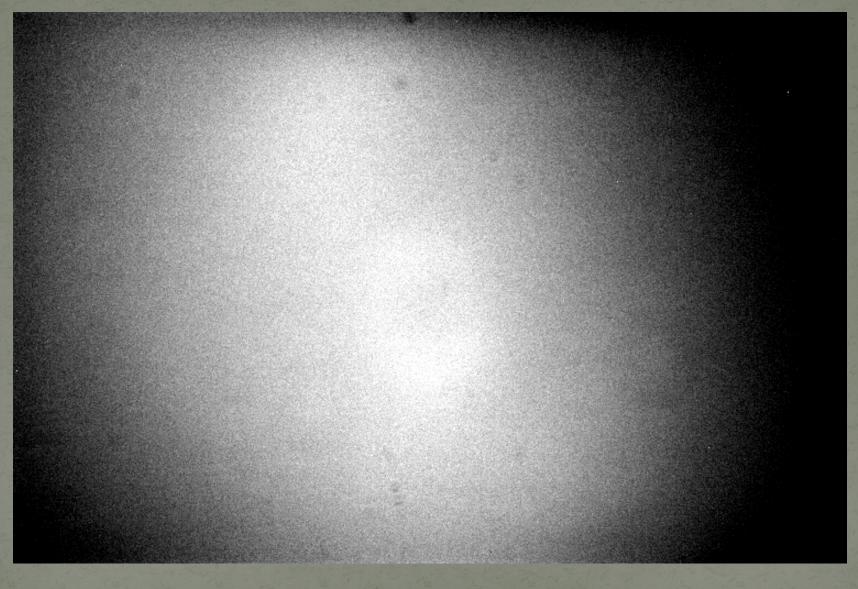
Calibration cont'd

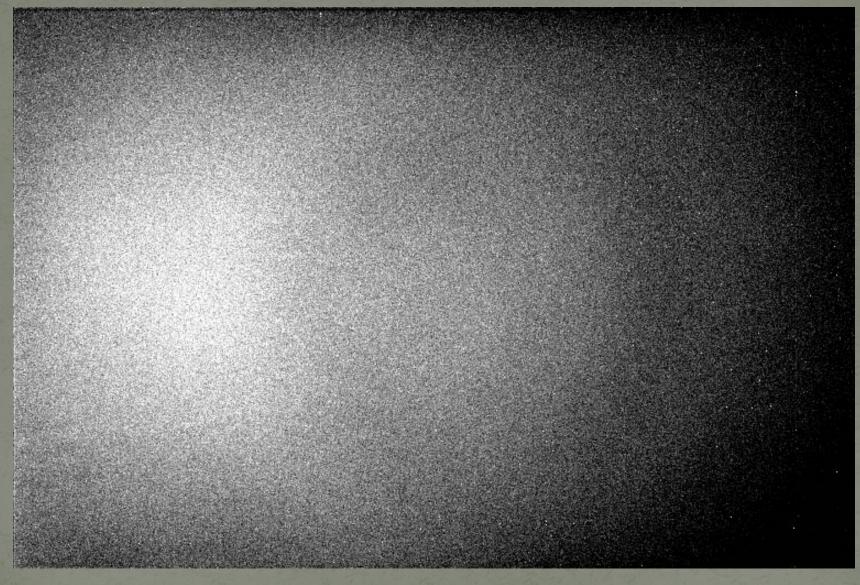
- What it won't do !!!
 - Correct for ...
 - tracking errors
 - field rotation
 - poor focus
 - field curvature
 - Coma
- Garbage in ... garbage out!

WHAT'S REQUIRED?

- How is it done?
 - subtract images of the optical and electronic/thermal
 « noise » from the captured astronomical frame
 - Subtract Flat frames to remove optical "noise"
 - Subtract Dark frames to remove electronic/thermal noise

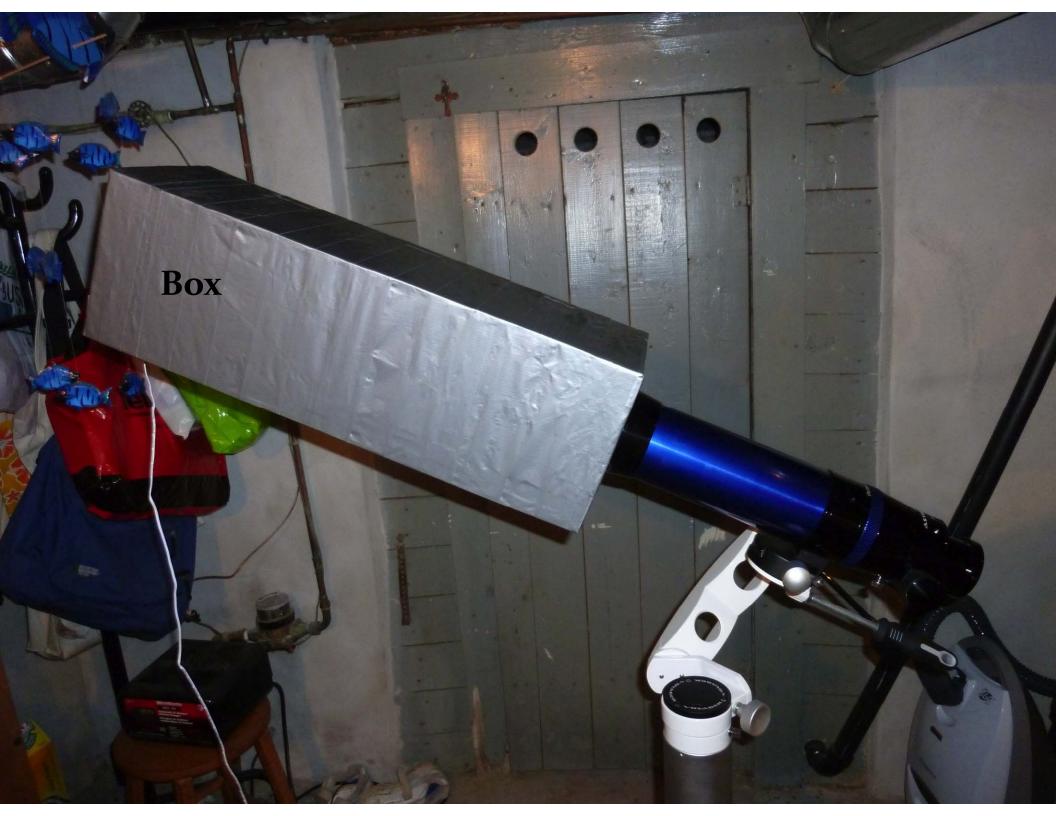
- An image of light passing through the telescope with no « object »
 - Use exactly the same telescope set-up as for the « real » frame capture
 - 1° Technical requirement: 33-50% pixel « saturation »





- An image of light passing through the telescope with no « object »
 - Requires even, diffuse illumination of the telescope/camera system
 - T-shirt method
 - Light box method
 - Flat panel method
 - 2° Technical requirement : minimum 2 seconds exposure



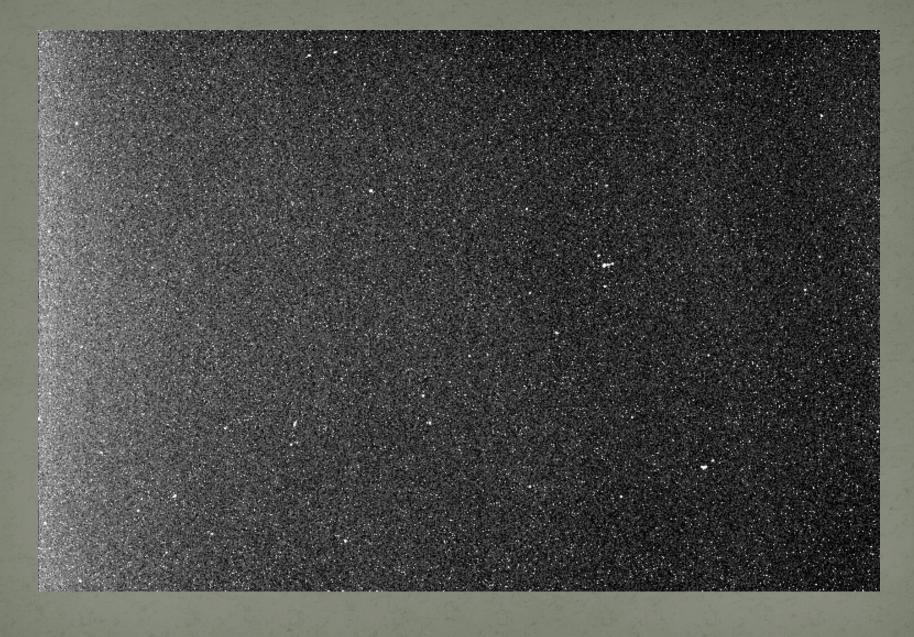




DARK FRAME

- An image of « bright » pixels when no light is falling onto the camera chip
 - Must use same chip temperature and exposure time as will be used during astro frame capture
 - Hardware set-up and camera orientation are not constrained
 - Frames valid for up to 3-6 months

Dark Frame



DARK FRAME cont'd

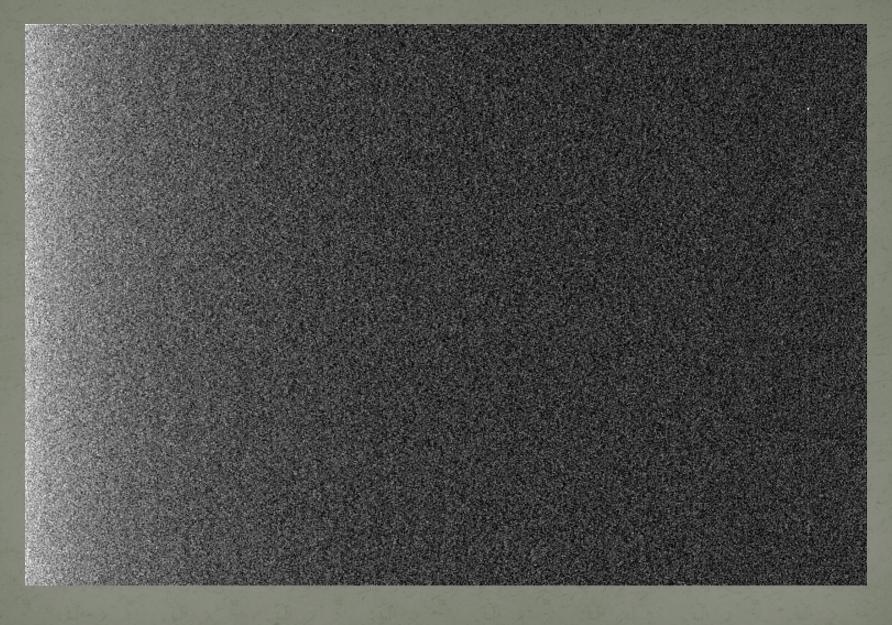
• An image of « bright » pixels when no light is falling onto the camera chip

• Includes "Darks for Flats" !!

BIAS FRAME

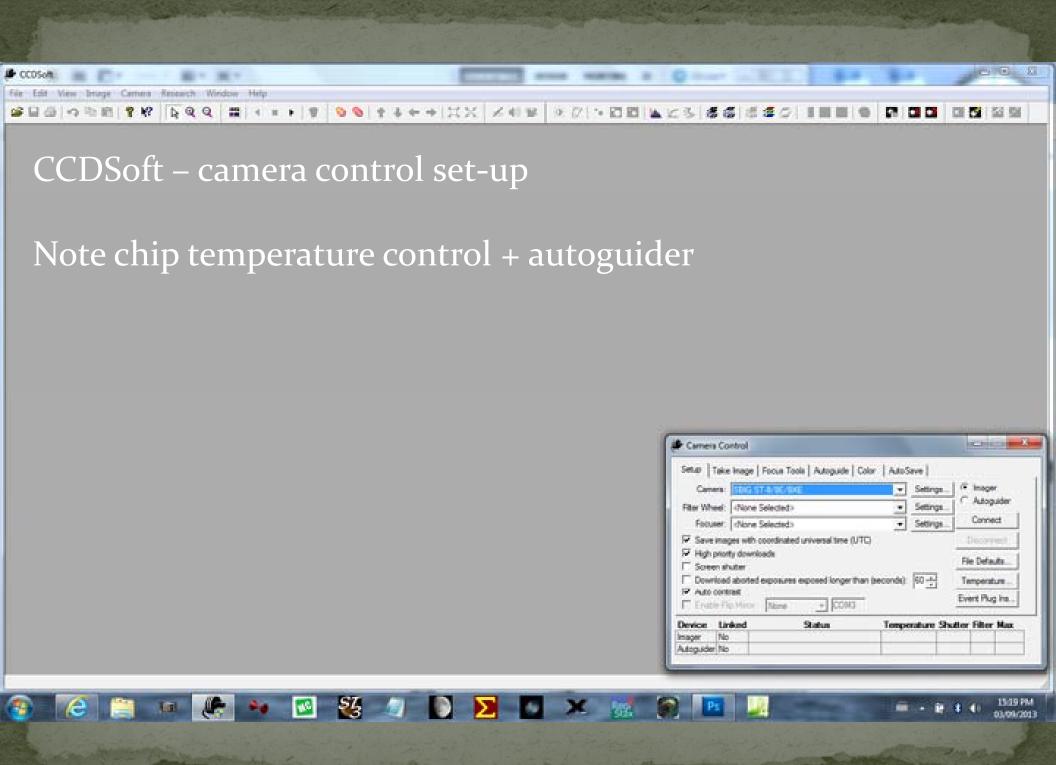
- A dark frame using the minimum exposure time available for your camera
 - Image of the fundamental electronic/thermal behaviour of the pixels
 - Bias frames used when astro capture exposure and/or temperature ≠ dark frames

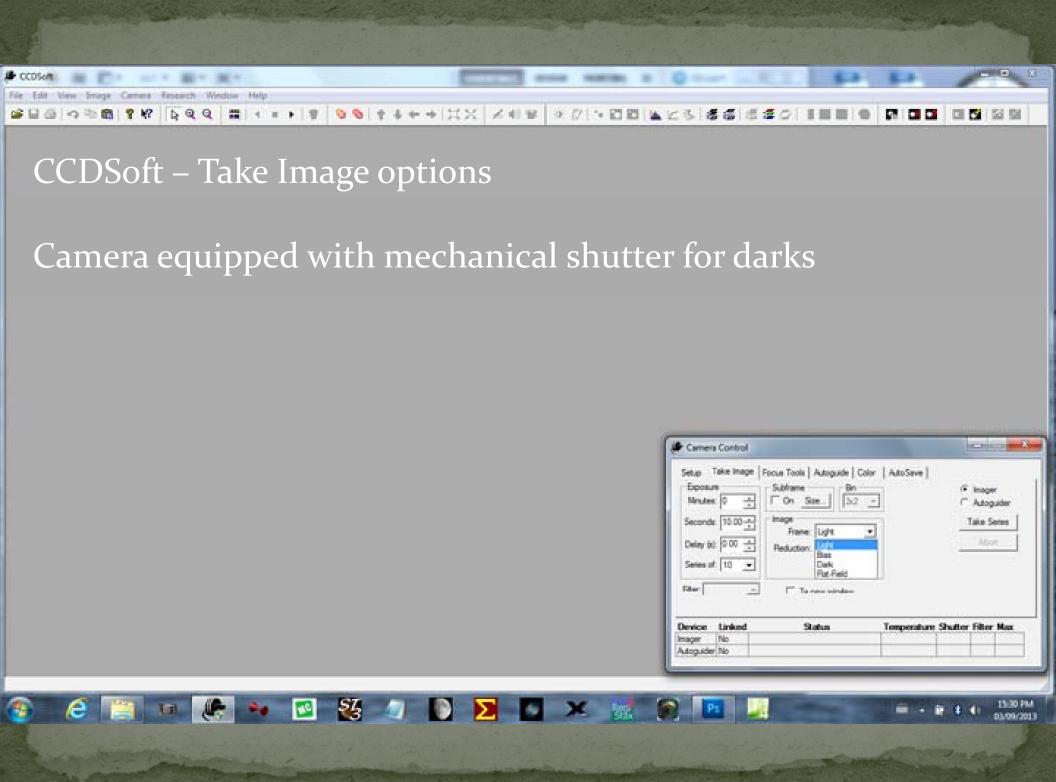
Bias Frame

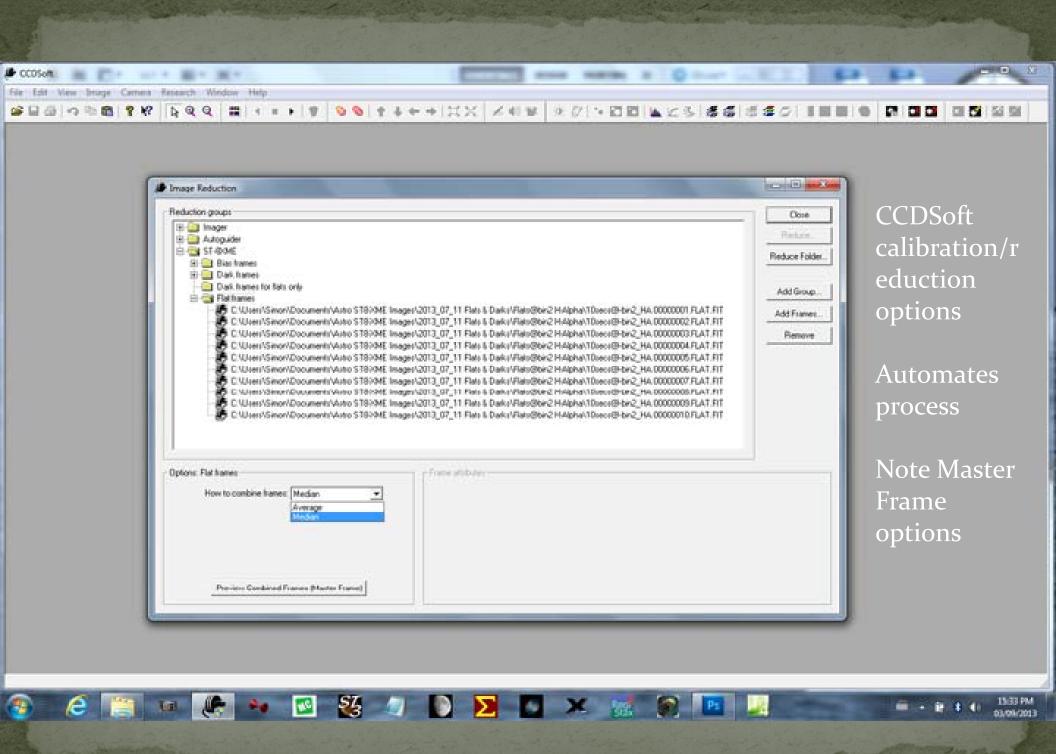


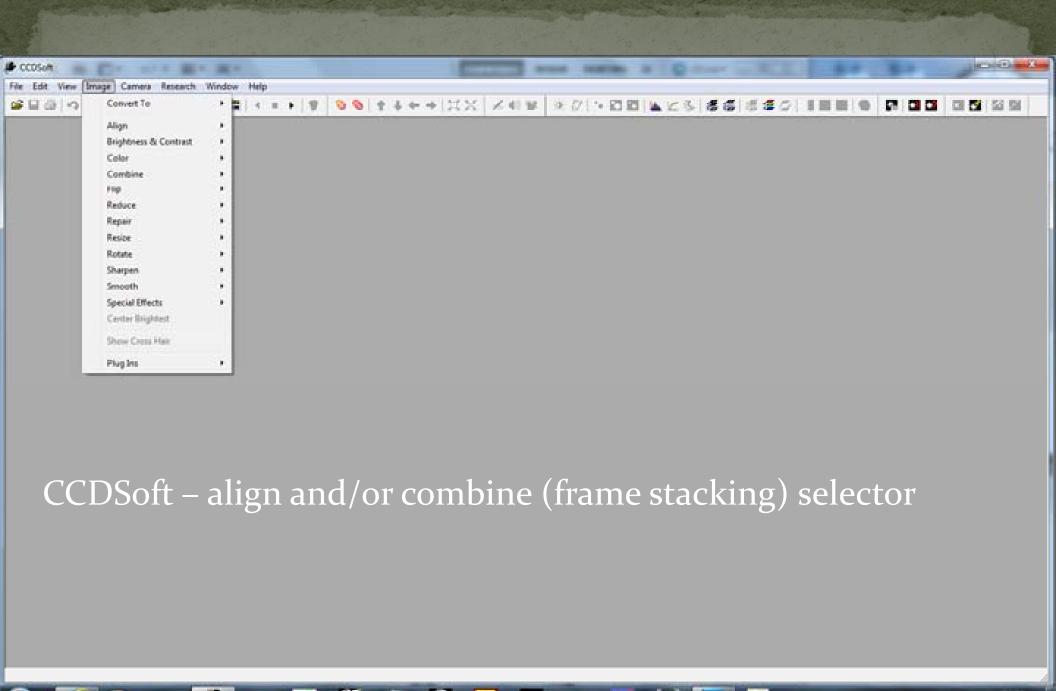
CALIBRATION - WORKED EXAMPLE

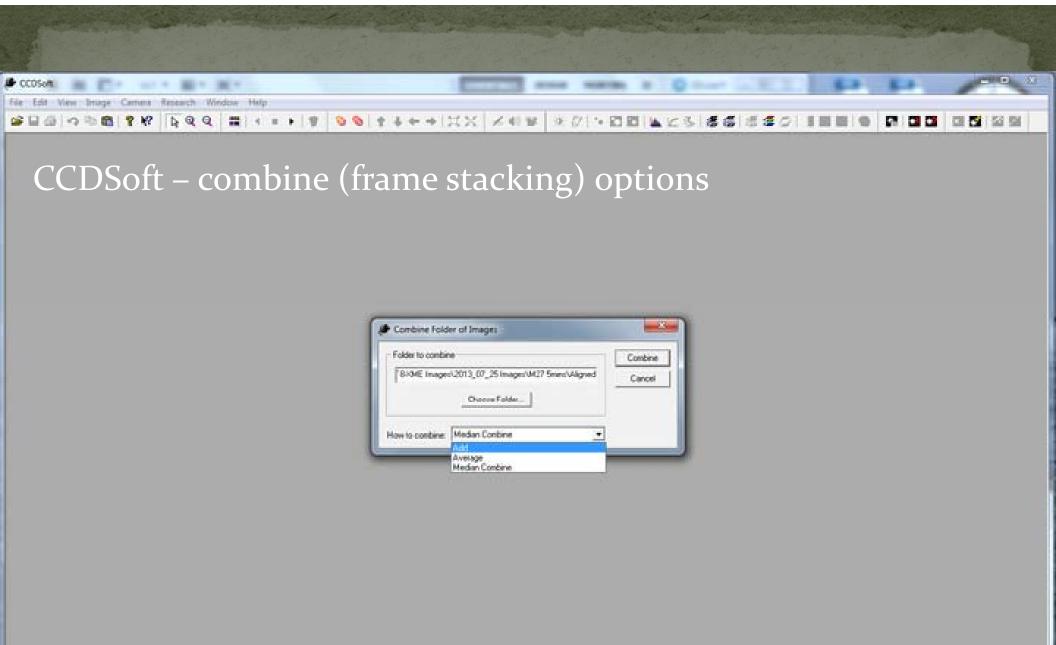
- In relation to CCD cameras ...
 - Abundant comprehensive software available ... we'll look at CCDSoft v5
 - CCD cameras are designed to make capture and calibration easy ... I use an SBIG ST8-XME
 - We'll now look at both these aspects ... in B&W : working in colour can require 3 times more work













HURDLES IN VIDEO CALIBRATION

- Camera ...
 - Estimating exposure i.e. % pixel saturation for Flat frames
 - Covering the scope for Dark and Bias frames
 - Determining the T°C of the camera chip for Dark and Bias frames

HURDLES IN VIDEO CALIBRATION

- Camera cont'd ...
 - 100% duty cycle = more thermal noise
 - High gain = more electronic noise and amp glow
 - Data fidelity limits « correction » potential

HURDLES IN VIDEO CALIBRATION

- Software...
 - No comprehensive package specifically for astro Video control + image frame calibration
 - Free- and Share-Ware packages, plus relatively inexpensive commercial software
 - It's up to you to find the appropriate packages for the calibration workflow ...
 - It's up to you to perform the work-arounds ...

HOWEVER ... if you're lucky!

- I captured 15 x 120 sec sub exposures and stacked them using Maxim DL along with my typical work flow in PSCS3 for the majority of the post processing
- Capturing the data is so easy and fast with a MC
 Xtreme and the post processing was about 15 minutes tops
- There was no image calibration involved



DSLR CALIBRATION

- DSLRs, with their larger size imaging chip, are usually used in wide field imaging
- Flats are especially important because the size of the imaging chip is large
- This will almost always result in uneven field illumination appearing as vignetting
- See also Dark & Bias frames (below)

DSLR CALIBRATION cont'd

- Other inherent issues
 - Mirrors
 - Sensor Cleaning
 - Noise Mapping
 - Noise Reduction

DSLR Camera Features

- Mirror Lock
 - BackYard EOS
- Noise Reduction
 - Turn it off!
- File Type
 - Start with JPEG!

DSLR Image Calibration

- Collecting DSLR Flats, Biases and Darks is tedious and confusing. Software can significantly reduce the effort.
- DeepSky Stacker (Freeware) organizes Lights Darks,
 Flats and Bias Frames and automatically applies them
- Astro Photography Tools
 (~\$20 CDN) automatically organizes and manages the capture of all calibration frames

DSLR FLATS

- DSLRs, with their larger size imaging chip, are usually used in wide field imaging
- Flats are especially important because the size of the imaging chip is large
- This will almost always result in uneven field illumination appearing as vignetting

- Dust motes on the chip are usually less obtrusive (especially in a DSLR with a Sensor Cleaning utility)
- Uneven light can result in either drop-off of details around the edges of the image, or an increase of noise in the center if the light drop-off at the edges is corrected.
- Amp glow is rarely a problem, however it may become evident in very long exposures (<10 minutes).
- Ideally Flats should be taken on a 1:1 ratio with the lights.
 Good practice is to have at least 30 Flats to average out.

- Flats need to be taken with the camera in the identical orientation and focus as the light frames
- Leave the camera untouched on the telescope
 - Take Flats either indoors with light screen

or

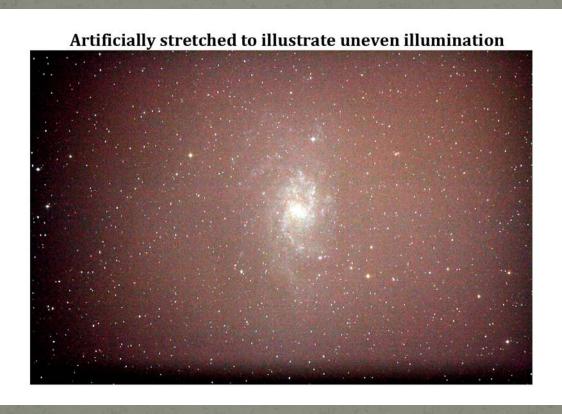
- Leave camera/telescope on the mount
 - Take Flats the next morning against a clear blue sky opposite the sun
- Set camera in AV mode
 - Let camera select the correct exposure
 - Take note of the shutter speed/exposure time
 - Shoot as many Flats as Light frames

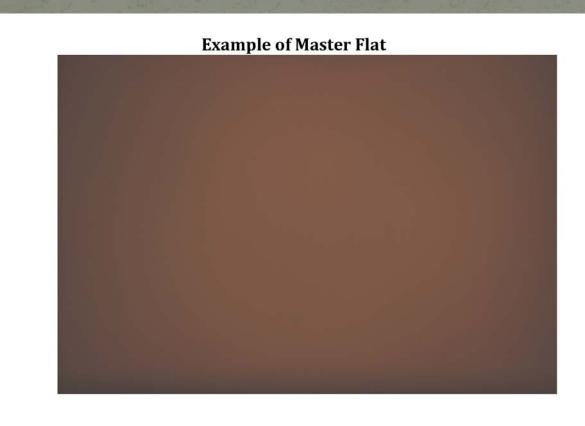
DSLR FLATS

Original Light Frame: (M33, 6in f/5 astograph, Canon D1000



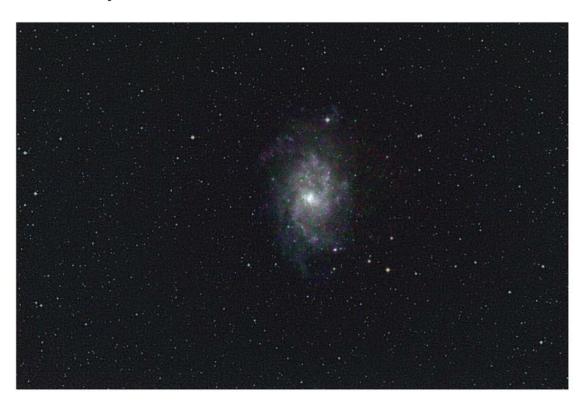
180 seconds at ISO 800 - single frame)

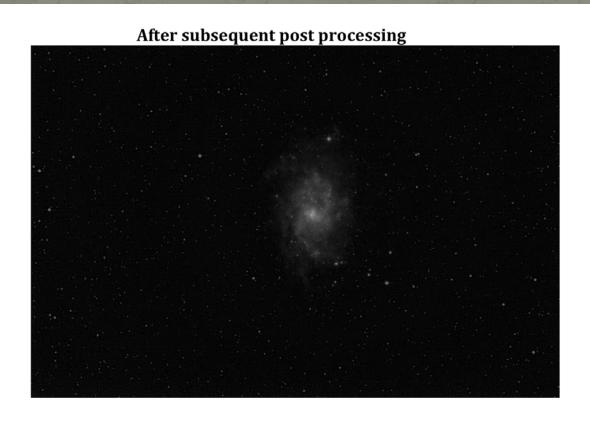






Artificially stretched to illustrate illumination correction

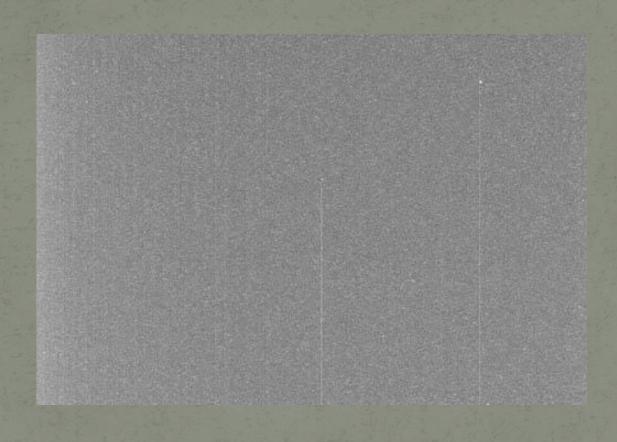




DSLR DARKS

- Dark frames will record hot pixels (red spots)
 and inherent noise (typically seen as horizontal
 or vertical banding)
- Removing these errors greatly improves the final image and is a superior method than dealing with hot pixels and banding in post-processing.

DSLR DARKS



DSLR DARKS cnt'd

- Collect your Dark frames at the same time as your Light frames
- They must be taken at the same camera settings and temperature as the lights, BUT with the telescope cap on
- Ideally, shoot at a 1:1 ratio with the Light frames

DSLR BIAS

- Bias Error is when the CMOS or CCD chip of the camera generates a signal that is created by the internal electronics of the camera just by reading the content/data
- Collect Bias frames the same time as Light frames
- Ideally, shoot at a 1:1 ratio with the Light frames
- Set the camera in Manual mode and dial to the fastest shutter speed available